

# Maxent Example - Self-Energy of a Metal

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## Abstract

This document is a tutorial on the use of Maxent, a program for doing analytical continuation using the maximum entropy method. It will explain how to provide the program with the proper parameter file, data format for particle-hole symmetric data of a self-energy Matsubara space, and understand the output. Included as a supplement is the corresponding interacting Green's function input data. The data provided is of an interacting Hubbard model with  $U = 1, \beta = 2$  for 1 site at half filling. This program uses the ALPSCore libraries[1, 2].

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## 1 Introduction

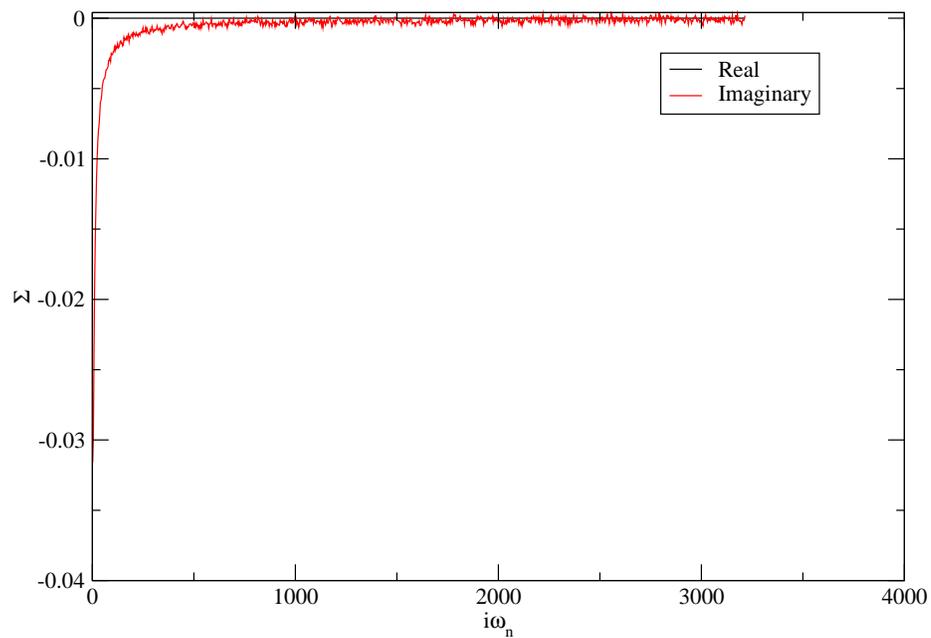
Using DMFT we can set  $U = 1, \beta = 2$  for a single site at half-filling to generate the interacting Hubbard model

$$H = - \sum_{\langle ij \rangle \sigma} t_{ij} \left( c_{i\sigma}^\dagger c_{j\sigma} + c_{j\sigma}^\dagger c_{i\sigma} \right) + U \sum_i n_{i\uparrow} n_{i\downarrow},$$

where  $t_{ij} = t = 1$ . This produces a self-energy output in Matsubara space:

### U=1 Self-Energy from DMFT

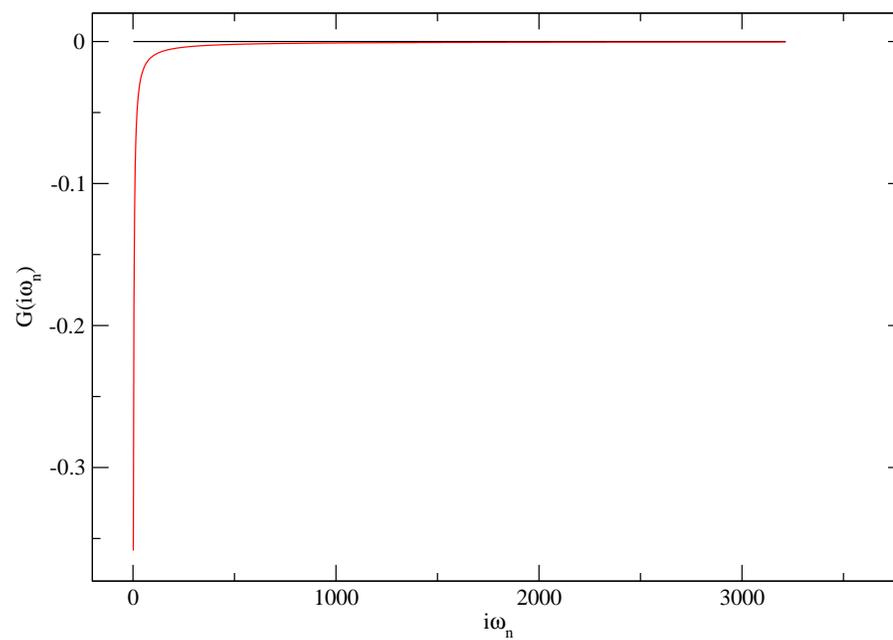
$\beta=2$



Measured separately is the Green's function:

### U=1 Green's Function

$\beta=2$



While these functions are on the imaginary axis, we wish to analytically continue them to the real axis. Mathematically, this is equivalent to finding the spectral function  $A(\omega) = \text{Im}[\Sigma(\omega)]$  such that

$$G(X) = -\frac{1}{\pi} \int_{-\infty}^{\infty} K(X, \omega) A(\omega) d\omega,$$

with the kernel  $K$  written as  $K(i\omega_n, \omega) = \frac{1}{i\omega_n - \omega}$  or  $K(\tau, \omega) = \frac{-e^{-\tau\omega}}{1 + e^{-\omega\beta}}$  for this data. The Maxent procedure is one such method of determining  $A(\omega)$ .

## 1.1 Normalization

Note that unlike a Green's function, the normalization (aka high frequency term) is not 1. For a self energy:

$$\Sigma = \Sigma_0 + \frac{\Sigma_1}{i\omega_n} + \dots$$

where  $\Sigma_0 = Un$  (known as the Hartree term) and  $\Sigma_1 = U^2n(1-n)$ [3]. This normalization is important to Maxent, as it is assumed to be the Green's function normalization of 1. For  $U = 1$  at half-filling ( $n = 0.5$ ), there is a normalization of 0.25.

## 1.2 Errors

In this case, DMFT self-energies are lacking error bars. Because errors go like  $\frac{1}{\sqrt{N}}$  where  $N$  is the number of samples, we can instead use this as an order of magnitude estimate for the error. This data had 4296900 sample iterations, and we therefore choose an estimated error  $\sigma = 0.0005$ .

## 2 File Structure

We've included several files that will be used to generate the remainder of this document:

| <b>Filenames and Descriptions</b>   |                         |  |
|---|-------------------------|--|
| <ul style="list-style-type: none"> <li>• <math>Selfenergy = \Sigma(i\omega_n)</math> data values               <ul style="list-style-type: none"> <li>– column format: <math>i\omega_n</math> <math>Re[\Sigma(i\omega_n)]</math> <math>Im[G(i\omega_n)]</math></li> </ul> </li> </ul> | Input file - Selfenergy |  |
| <hr/>   |                         |  |
| <pre>1.5707963267949 5.6703744085571e-16 -0.03158953764285 4.7123889803847 3.1444695800992e-16 -0.029847318775471 7.8539816339745 9.4394785839476e-16 -0.023814398782168</pre>  |                         |  |
| <hr/>   |                         |  |
| <ul style="list-style-type: none"> <li>• <math>Selfin = \Sigma(i\omega_n)</math> input format for Maxent               <ul style="list-style-type: none"> <li>– column format: <math>i\omega_n</math> <math>Im[G(i\omega_n)]</math> <math>\sigma_{I,n}</math></li> </ul> </li> </ul>  | Input file - Selfin     |  |
| <hr/>   |                         |  |
| <pre>1.5707963267949 -0.03158953764285 0.0005 4.7123889803847 -0.029847318775471 0.0005 7.8539816339745 -0.023814398782168 0.0005</pre>   |                         |  |
| <hr/>   |                         |  |
| <ul style="list-style-type: none"> <li>• <math>G\_im = Im[G(i\omega_n)]</math>, also input for Maxent               <ul style="list-style-type: none"> <li>– column format: <math>i\omega_n</math> <math>Im[G(i\omega_n)]</math> <math>\sigma_{I,n}</math></li> </ul> </li> </ul>     | Input file - G_im       |  |
| <hr/>   |                         |  |
| <pre>1.5707963267949 -0.3585154015692 0.0005 4.7123889803847 -0.18353508163336 0.0005 7.8539816339745 -0.11976613812711 0.0005</pre>  |                         |  |
| <hr/>   |                         |  |
| <ul style="list-style-type: none"> <li>• <math>G\_re = Re[G(i\omega_n)]</math> <ul style="list-style-type: none"> <li>– column format: <math>i\omega_n</math> <math>Re[G(i\omega_n)]</math> <math>\sigma_{R,n}</math></li> </ul> </li> </ul>  | Input file - G_re       |  |
| <hr/>   |                         |  |
| <pre>1.5707963267949 -2.0386388974517e-17 0.0005 4.7123889803847 -4.6688456052657e-18 0.0005 7.8539816339745 -3.4351421175898e-18 0.0005</pre>  |                         |  |
| <hr/>   |                         |  |

### 3 Using Maxent

These files are easily used with Maxent. Here is the frequency space input:

---

Param File in.param

---

```

BETA=2                #inverse temperature
OMEGA_MAX=25          #the spectral function is wider than omega=10
NDAT=1024              #num of data points
NFREQ=1000            #num of output frequencies
DATASPACE=frequency   #G(iω)
KERNEL=fermionic      #fermionic/bosonic values
FREQUENCY_GRID=Quadratic #this grid is better for features away from 0
PARTICLE_HOLE_SYMMETRY=1 #generated at half-filling
DATA="Selfin"         #location of data file
SELF=1                #this will output Σ(ω) rather than A(ω)
NORM=0.25             #self energy norm = U2·n(1-n)

```

---

Maxent then produces the following output:

---

Maxent output

---

```

Using flat default model
using kernel fermionic in domain frequency with ph symmetry
The high frequency limit is not 1!: 2.30495 Check norm?
Kernel is set up
# 0      4108.32
# 1      1668.62
# 2      513.561
# 3      126.175
# 4      25.6169
# 5      4.3795
# 6      0.635099
# 7      0.0784661
minimal chi2: 0.098459
WARNING: Redefinition of parameter NORM: Input (and output) data are assumed to be
        normalized to NORM.
alpha it: 0      Q = 0.5chi^2-\alpha*entropy: 527.241      norm: 1.16962
alpha it: 1      Q = 0.5chi^2-\alpha*entropy: 481.411      norm: 1.17798
alpha it: 2      Q = 0.5chi^2-\alpha*entropy: 443.307      norm: 1.18585
alpha it: 3      Q = 0.5chi^2-\alpha*entropy: 409.139      norm: 1.19303
alpha it: 4      Q = 0.5chi^2-\alpha*entropy: 378.828      norm: 1.19943
alpha it: 5      Q = 0.5chi^2-\alpha*entropy: 352.116      norm: 1.20499
alpha it: 6      Q = 0.5chi^2-\alpha*entropy: 328.667      norm: 1.2097
alpha it: 7      Q = 0.5chi^2-\alpha*entropy: 308.1         norm: 1.21353
alpha it: 8      Q = 0.5chi^2-\alpha*entropy: 290.018      norm: 1.21649
alpha it: 9      Q = 0.5chi^2-\alpha*entropy: 274.026      norm: 1.21858
alpha it: 10     Q = 0.5chi^2-\alpha*entropy: 259.754      norm: 1.21982
alpha it: 11     Q = 0.5chi^2-\alpha*entropy: 246.864      norm: 1.22024
alpha it: 12     Q = 0.5chi^2-\alpha*entropy: 235.057      norm: 1.21985
alpha it: 13     Q = 0.5chi^2-\alpha*entropy: 224.078      norm: 1.21869
alpha it: 14     Q = 0.5chi^2-\alpha*entropy: 213.717      norm: 1.21679
alpha it: 15     Q = 0.5chi^2-\alpha*entropy: 203.807      norm: 1.21418
alpha it: 16     Q = 0.5chi^2-\alpha*entropy: 194.22        norm: 1.2109
alpha it: 17     Q = 0.5chi^2-\alpha*entropy: 184.869      norm: 1.20697
alpha it: 18     Q = 0.5chi^2-\alpha*entropy: 175.7         norm: 1.20245
alpha it: 19     Q = 0.5chi^2-\alpha*entropy: 166.688      norm: 1.19737
alpha it: 20     Q = 0.5chi^2-\alpha*entropy: 157.834      norm: 1.19178
alpha it: 21     Q = 0.5chi^2-\alpha*entropy: 149.256      norm: 1.18533
alpha it: 22     Q = 0.5chi^2-\alpha*entropy: 140.793      norm: 1.17884

```

```

alpha it: 23    Q = 0.5chi^2-\alpha*entropy: 132.591    norm: 1.17201
alpha it: 24    Q = 0.5chi^2-\alpha*entropy: 124.7      norm: 1.1649
alpha it: 25    Q = 0.5chi^2-\alpha*entropy: 117.171    norm: 1.15758
alpha it: 26    Q = 0.5chi^2-\alpha*entropy: 110.049    norm: 1.15012
alpha it: 27    Q = 0.5chi^2-\alpha*entropy: 103.374    norm: 1.14258
alpha it: 28    Q = 0.5chi^2-\alpha*entropy: 97.1752     norm: 1.13502
alpha it: 29    Q = 0.5chi^2-\alpha*entropy: 91.4723     norm: 1.1275
alpha it: 30    Q = 0.5chi^2-\alpha*entropy: 86.2731     norm: 1.12008
alpha it: 31    Q = 0.5chi^2-\alpha*entropy: 81.5754     norm: 1.11281
alpha it: 32    Q = 0.5chi^2-\alpha*entropy: 77.3674     norm: 1.10573
alpha it: 33    Q = 0.5chi^2-\alpha*entropy: 73.6294     norm: 1.09887
alpha it: 34    Q = 0.5chi^2-\alpha*entropy: 70.3353     norm: 1.09227
alpha it: 35    Q = 0.5chi^2-\alpha*entropy: 67.4542     norm: 1.08595
alpha it: 36    Q = 0.5chi^2-\alpha*entropy: 64.9523     norm: 1.07993
alpha it: 37    Q = 0.5chi^2-\alpha*entropy: 62.7942     norm: 1.07421
alpha it: 38    Q = 0.5chi^2-\alpha*entropy: 60.944      norm: 1.0688
alpha it: 39    Q = 0.5chi^2-\alpha*entropy: 59.3668     norm: 1.0637
alpha it: 40    Q = 0.5chi^2-\alpha*entropy: 58.029      norm: 1.05891
alpha it: 41    Q = 0.5chi^2-\alpha*entropy: 56.8992     norm: 1.05441
alpha it: 42    Q = 0.5chi^2-\alpha*entropy: 55.9485     norm: 1.05021
alpha it: 43    Q = 0.5chi^2-\alpha*entropy: 55.1507     norm: 1.04629
alpha it: 44    Q = 0.5chi^2-\alpha*entropy: 54.4826     norm: 1.04263
alpha it: 45    Q = 0.5chi^2-\alpha*entropy: 53.9235     norm: 1.03923
alpha it: 46    Q = 0.5chi^2-\alpha*entropy: 53.4556     norm: 1.03606
alpha it: 47    Q = 0.5chi^2-\alpha*entropy: 53.0636     norm: 1.0331
alpha it: 48    Q = 0.5chi^2-\alpha*entropy: 52.7344     norm: 1.03035
alpha it: 49    Q = 0.5chi^2-\alpha*entropy: 52.457      norm: 1.02779
alpha it: 50    Q = 0.5chi^2-\alpha*entropy: 52.2225     norm: 1.02541
alpha it: 51    Q = 0.5chi^2-\alpha*entropy: 52.0232     norm: 1.02319
alpha it: 52    Q = 0.5chi^2-\alpha*entropy: 51.8531     norm: 1.02112
alpha it: 53    Q = 0.5chi^2-\alpha*entropy: 51.7072     norm: 1.01919
alpha it: 54    Q = 0.5chi^2-\alpha*entropy: 51.5814     norm: 1.01739
alpha it: 55    Q = 0.5chi^2-\alpha*entropy: 51.4727     norm: 1.01571
alpha it: 56    Q = 0.5chi^2-\alpha*entropy: 51.3782     norm: 1.01415
alpha it: 57    Q = 0.5chi^2-\alpha*entropy: 51.2961     norm: 1.01269
alpha it: 58    Q = 0.5chi^2-\alpha*entropy: 51.2245     norm: 1.01133
alpha it: 59    Q = 0.5chi^2-\alpha*entropy: 51.162      norm: 1.01006

```

Ng: 5.10747

posterior probability of the default model: 1.33065e-33

| spectra | max backcont diff | chi^2 value |
|---------|-------------------|-------------|
| chispec | 0.040974          | 1053.08     |
| avspec  | 0.00203636        | 106.161     |
| maxspec | 0.00195274        | 105.411     |

### 3.1 Output Guide

```

Using flat default model
using kernel fermionic in domain frequency with ph symmetry
The high frequency limit is not 1!: 2.30495 Check norm?
Kernel is set up

```

These are the setup messages, confirming your input choices. There is a warning for the high frequency limit, but because our data is very noisy it can be ignored. If this limit was significantly off from 1, then your input NORM should be confirmed. In this case, the last few data points are noisy, leading to an inaccurate high frequency limit warning.

```
# 0 4108.32
# 1 1668.62
# 2 513.561
# 3 126.175
# 4 25.6169
# 5 4.3795
# 6 0.635099
# 7 0.0784661
minimal chi2: 0.098459
```

These represent the eigenvalues that are above precision after the single value decomposition (SVD). The last line represents the smallest  $\chi^2$  value the program thinks it will achieve. If this is  $\gg 1$  there may be something wrong with your input or it is very noisy

```
...
alpha it: 2 Q = 0.5chi^2-\alpha*entropy: 443.307 norm: 1.18585
alpha it: 3 Q = 0.5chi^2-\alpha*entropy: 409.139 norm: 1.19303
...
```

The root finding procedure will print the iterations through  $\alpha$  values in the range given by the parameters (default: 60 values  $\in [0.01, 20]$ ) If the first two or three do not minimize properly that is ok, as long as the rest continue normally. Notice that the norm stays  $\approx 1$  for all iterations

```
Ng: 5.10747
chi2 max: 105.411
posterior probability of the default model: 1.33065e-33
```

This is posted after completing all  $\alpha$  values and root finding. Ng represents the number of “good input points,” chi2 max is the maximum value of  $\chi^2$  in the  $\alpha$  iterations, and the last line is the probability that the default model is the correct representation of the spectral function. Note that that posterior probability has no known normalization.

| spectra | max backcont diff | chi^2 value |
|---------|-------------------|-------------|
| =====   | =====             | =====       |
| chispec | 0.040974          | 1053.08     |
| avspec  | 0.00203636        | 106.161     |
| maxspec | 0.00195274        | 105.411     |

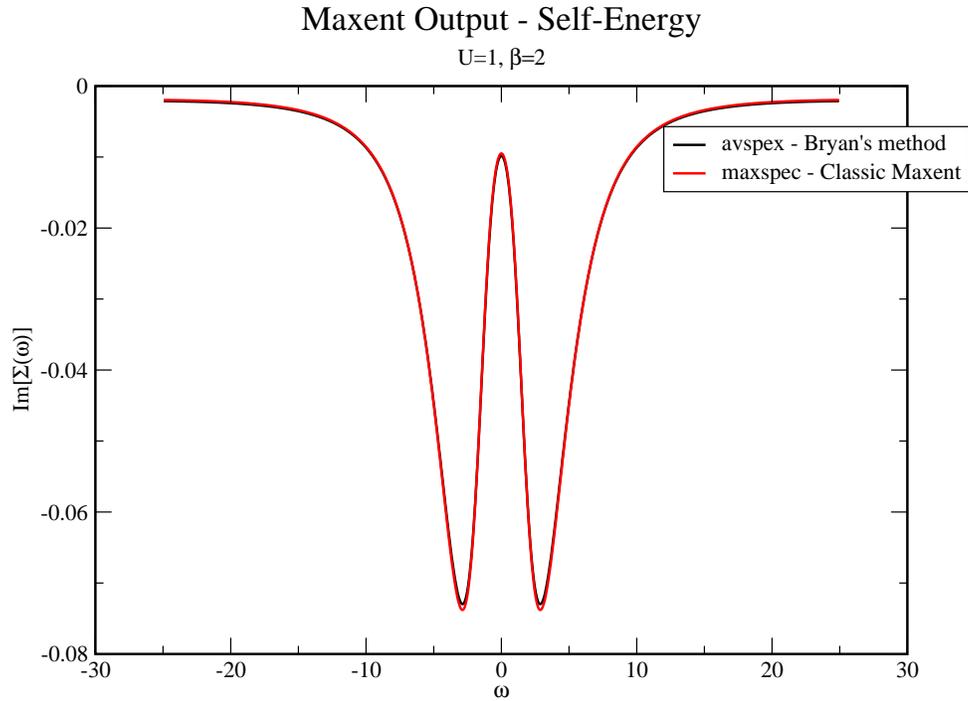
By default, maxent will back-continue, or continue back to the imaginary axis, the spectral function maxent output. Here two useful values are shown, the maximum difference between any of the back-continued points and input data, as well as the  $\chi^2$  value.

---

If text output is on, Maxent produces 13 files:

|                           |   |
|---------------------------|---|
| name.out.avspec.dat       | “Spectral function” using Bayesian Averaging - <b>Bryan’s method</b>                            |
| name.out.avspec_back.dat  | The avspec spectrum continued back to the imaginary axis  |
| name.out.avspec_self.dat  | $\text{Im}[\Sigma(\omega)]$ with the proper sign and normalization; using <b>Bryan’s method</b> |
| name.out.chi2.dat         | Estimated $\chi^2$ for each $\alpha$ value solution   |
| name.out.chispec.dat      | “Spectral function” satisfying the best $\chi^2$ - <b>historic Maxent</b>                       |
| name.out.chispec_back.dat | The chispec spectrum continued back to the imaginary axis                                       |
| name.out.fits.dat         | Fits of each $\alpha$ value, see comments in file   |
| name.out.maxspec.dat      | “Spectral function” with the highest probability - <b>classic Maxent</b>                        |
| name.out.maxspec_back.dat | The maxspec spectrum continued back to the imaginary axis                                       |
| name.out.maxspec_self.dat | $\text{Im}[\Sigma(\omega)]$ with the proper sign and normalization; using <b>classic Maxent</b> |
| name.out.out.h5           | All output data in the hdf5 format  |
| name.out.prob.dat         | The posterior probability of each $\alpha$ value  |
| name.out.spex.dat         | All spectral functions produced; one for each $\alpha$  |

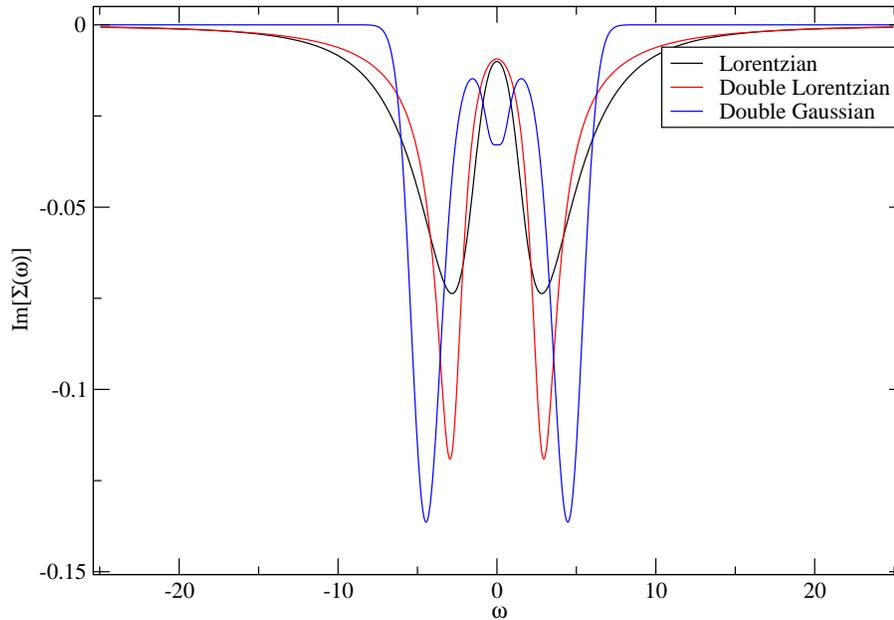
Because this is a self-energy, Maxent treats the input as a Green's function and finds a spectral function associated with it, but the spectral function output itself is meaningless. In our example here are the self-energy outputs with a flat default model:



## 4 Fine-Tuning Output

Different default models shouldn't change the results much, but sometimes end up doing so. Here are a variety of models from the above example:

## Maxent With Various Default Models Using Bryan's Method



With  $\sigma = 1, \Gamma = 1, \mu(\text{shift}) = 2.8$ . The default model gives a spectral function most similar to a double Lorentzian. When provided with a double Gaussian, Maxent attempts to fit the center peak, but is overcome with the entropy from the Gaussian model underneath.

## References

- [1] B Bauer et al. The ALPS project release 2.0: open source software for strongly correlated systems. *Journal of Statistical Mechanics: Theory and Experiment*, 2011(05):P05001, 2011. doi:10.1088/1742-5468/2011/05/P05001.
- [2] Alexander Gaenko, Emanuel Gull, Andrey E. Antipov, Lukas Gamper, Gabriele Carcassi, Joe Paki, Ryan Levy, Michele Dolfi, Jonas Greitemann, and James P. F. LeBlanc. ALPSCore: Version 0.5.4. April 2016. doi:10.5281/zenodo.50203.
- [3] Xin Wang, Emanuel Gull, Luca de' Medici, Massimo Capone, and Andrew J. Millis. Antiferromagnetism and the gap of a mott insulator: Results from analytic continuation of the self-energy. *Phys. Rev. B*, 80:045101, Jul 2009. URL: <http://link.aps.org/doi/10.1103/PhysRevB.80.045101>, doi:10.1103/PhysRevB.80.045101.